Lahroud a Paleo-Tethys Remnant in northwestern Iran, Implications to: Geochemistry, Radioisotope, Geochronology and Tectonic setting Shohreh Hassanpour

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Abstract

The Lahroud Ophiolite in northwestern Iran contains extensive zones of Paleozoic ophiolite as remnants of the Paleo-Tethys oceanic crust. The principal rock units are gabbro overlain by pillow basalt which is intruded by granites, interbedded with pelagic sedimentary units, which includes radiolarian cherts. Geochemistry and radioisotope studies, supported by Nd, Sm, Sr, and Pb isotope data indicate that the Lahroud ophiolite has been raised from a within plate basaltic mantle source. Isotope studies show that the basalts are derived from Indian type Oceanic mantle sources. Radiogenic data indicates the involvement of subduction related terrigenous materials in the source magma. All rocks are geochemically cogenetic and were generated by fractionation of a melt with a composition of average E-MORB with a calcalkaline signature.

Two ⁴⁰Ar/³⁹Ar ages of 343±3 and 187.7±7.7 Ma for the muscovite minerals and glasses respectively, suggest that metamorphic and basaltic rocks were formed during the Late Paleozoic to Early Jurassic, respectively.

Microfossil studies show the presence of Paleozoic biostratigraphy. The crystallization process and rifting into the oceanic crust in the Lahroud ophiolite probably began in the Carboniferous, with volcanic activity continuing during the late Triassic.

Keywords: Ophiolite, Lahroud, Paleotethys, Azerbaijan, Iran

Introduction

Oceanic lithosphere fragments remained as ophiolites and were accreted through orogenic phases. The present day locations of the ophiolites represent a suture zone, and are important markers in the reconstruction of the tectonic history (Dilek, 2003). A complete ophiolite typically comprises, from top to bottom: pillow lava, sheeted dykes, and cumulate mafic to ultramafic units (Anonymous, 1972), however not all preserved ophiolites contain a complete sequence. Commonly, ophiolites are strongly deformed and metamorphosed during uplift. Iranian ophiolites and their locations are illustrated in Fig. 1 and have been classified in several ways by different authors. Arvin and Robinson (1994) divided Iranian ophiolites into two groups based on their ages: Paleozoic ophiolites in the north of Iran (less common) and Mesozoic ophiolites (more common) in the south of Iran, near the main central Iranian block ophiolites (e.g., Shafaii Moghadam and Stern, 2011). Recently, Iranian ophiolites have

been classified into four groups based on their location and geotectonic setting: (i) northern ophiolites along the Alborz range (ii) Zagros ophiolites as a Suture Zone (iii) unfragmented ophiolites in the Makran zone; and (iv) Central Iranian Microcontinent ophiolites and colored melanges (CIM) (McCall, 1997; Stocklin, 1974; Takin, 1972; Fig. 1; Table 1).



Figure1. Simplified regional geological map of Iran emphasizing the main ophiolitic belts (thick dashed lines) (modified from Shafaii Moghadam et al., 2014).

References	Emplacement age (Ma)	Formation age (Ma)	Ophiolite Complex	Region
Ghazi et al. (1997a)	106-110 (Ar/Ar)	154-159 (Ar/Ar)	Khoy	N-NW
Hassanipak and Ghazi (2000)	143 (Muscovite)			
Alavi (1996)	(?)	Pre Jurassic	Rasht	Zagros
Lippard et al. (1986)	Post-Campanian	95-98 (Ar/Ar)	Kermanshah	
Ghazi and Hassanipak (1999)	Pre- Maastrichtian	(?)		
Delaloye and Desmons (1980)	(?)	86.3?	Sahneh	
Ghazi et al. (1999), Lanphere and		93-95 Ma		
Pamic (1983), Sarkarinejad (1994)	89 Ma (Ar/Ar)	(Ar/Ar)	Neyriz	
	Post-Campanian	96-98 Ma (Ar/Ar)		
	Pre- Maastrichtian			
Lippard et al. (1986), Sarkarinejad (1994), Hassanipak and Ghazi (2000)	Pre-Paleocene	93-95 Ma (Ar/Ar)	Naein	South- Central
Campbell et al. (2000), Aftabi and Fathi (1992), Tucker et al.	Pre-Paleocene	93 Ma (Ar/Ar)	Shar-e-Babak	
	Pre-Paleocene	Upper Cretaceous	Baft-Esphandagheh	
Ghazi et al. (1997b), Hassanipak				
et al. (1996), Desmons and Beccaluva (1983), McCall (1997), Ghazi et al. (1999)	Early Paleocene	140-142 (Ar/Ar)	Band-e-Zeyarat	Makran
			Dar Anar	
McCall (1985),	Early Eocene	135 K/Ar)	Ganj	
McCall (1985),Desmons and Beccaluva (1983), McCall (1997), Hassanipak et al. (1996)	Early Paleocene	Late Jurassic-	Remeshk/Mokhtar	
		Early Cretaceous	Abad	
	Pre-Paleocene	Upper Crota accus	Iranshar	
	Pre-Paleocene	Upper Cretaceous Upper	Fanuj-makustan	
Desmons and Beccaluva (1983). Lippard et al. (1986), Delalove		Upper		East-
and Desmons (1980)	Pre Paleocene	Cretaceous	TchehelKureh	Northeast
Hassanipak et al. (2002), Lensch et al. (1977), Lensch (1980), Shojaat (1999)	Post Paleocene	52.4 (Ar/Ar)	sabzevar	
Ghazi et al. (1999), Alavi (1996), Alavi (1992), Hassanipak et al. (2002)	Pre Jurassic	277-281 (Ar/Ar)	Mashhad	
This study	Paleozoic	245-187 Ma (Ar/Ar)	Lahroud (Alahyarlou)	NW

Table 1: Distribution of Iran different ophiolites

The CIM comprises the Lut, Posht Badam, Tabas, and the Yazd ophiolites (Fig. 1; Table 1). Alavi (1991) classified Iranian ophiolites through observations of field relationships into three groups: (i) the Proterozoic, which is presented in the western edge of the CIM as an isolated outcrop; (ii) Pre-Jurassic, which is located in the Alborz E-W Range; and (iii) Post-Jurassic, which is the most abundant. Table 1 shows distributions, age, and tectonic settings of ophiolites in Iran (Ghazi et al., 2003). On Fig. 1, the location of Paleo-Tethys ophiolites are traced in an east- west trend. These ophiolitic slices in the north of Iran are comprised of



Darrehanjir, Mashhad, Fariman (in Kopet Dagh), Jandagh-Anarak, Takab, and Rasht ophiolites (see Fig. 1; Shafaii Moghadam et al., 2014).

Figure 2. Simplified local geological map of ophiolitic sequence in the Lahroud area (see Fig. 3) with sample locations (After Amidi et al., 1991).

Paleozoic ophiolites in the southwestern part of Asia are exposed in the Caucasus across Iran, Turkey (Kure mélange), Turkmenistan, Afghanistan, and Tibet (Shi et al., 2012; Meng et al., 2013). The distribution of Paleozoic ophiolites along the southern margin of Eurasia represents the closure of the Paleo-Tethys oceanic basin (e.g., Omrani et al., 2013; Zanchetta et al., 2013; Rolland et al., 2011; Dai et al., 2011; Bagheri and Stampfli, 2008; Stampfli and Borel, 2002). In the southern part of Eurasia, the Paleo-Tethys Ocean formed during the Upper Paleozoic. Closure of the Paleo-Tethys commenced in the Middle to Late Jurassic (the Eo-Cimmerian deformation (Zanchi et al., 2009; Boulin, 1988) resulting in the formation of modern day northern Iran and Afghanistan. Paleo-Tethys rifting produced alkalic-tholeiitic continental flood type basalts (Soltan-Meidan basalts), felsic to mafic plutonic rocks ("ages of ca. 460 Ma"; Shafaii Moghadam and Stern, 2014), dolomite, evaporites, and terrigenous sediments in the Ordovician (Ghelli) and Lower Devonian (Padeha) formations in the northern part of Iran (Aharipour et al., 2010; Stampfli, 1978). Studies showed that, following Ordovician to Silurian rifting, seafloor spreading of the Paleo-Tethys was diachronous, and

vary from Devonian to Permian. Prior to this study, the age of closure of the Paleo-Tethys in northwestern Iran was (in east-west trend) unknown. This paper presents the results of studies on the formation and evolution of the Lahroud (Allahyarlou) ophiolite in northwestern Iran, including petrographic studies, chemical and isotope analyzes, field observations and radiometric age dating. The results were used to determine the stratigraphy, igneous rock composition, and petrogenesis in order to discuss a tectonic environment and biostratigraphic setting of the ophiolite in the context of the closure of the Paleo-Tethys Ocean.

Regional geology

The Lahroud area in northwestern part of Iran is a relatively small very poorly known outcrop that is situated in the Alborz zone (Stocklin, 1968; Berberian and Amidi, 1981) and the southern part of the Caspian depression. The Lahroud ophiolite is a part of several ophiolite outcrops in northeastern Iran. All of some ophiolitic patches are distributed tectonically in the boundary between the CIM and the Alborz mountains in an east-west trend among Gondwana and Eurasia plates (Paleo-Tethys suture zone). This ophiolite is considered to represent a part of the northern branch of the Paleo-Tethys Ocean that opened and closed during the Paleozoic and Early Jurassic and proposed to be related to the ophiolites of the Lesser Caucasus.

In the Lahroud (Alahyarlou) area, three important geologic features in the region are distinguished: (a) Lahroud ophiolite sequence, which is composed of gabbro, ultramafic rocks, and unconformable overlain by (b) Upper Jurassic to Cretaceous pillow lava not from the ophiolite sequence, which overlies ophiolitic sequences; and (c) Pelagic limestones in some patches in the region and around the sequences (Figs. 2, 3). The Lahroud sequence comprises gneiss, amphibolite, chlorite schist, and meta-diabase that are associated with serpentinites, gabbro, dunite, tectonic breccia, and marbles. The complex has an ultramaficgabbroic unit that contains some sequences of gabbroic rock units that are preserved and overlain by sheeted diabase dykes. Other dykes which strike parallel to the ophiolite are also observed (Fig. 2). The Lahroud sequence is weakly serpentinized. The pelagic limestone strike is east-west. Pelagic limestones are pink and red in color with intercalation of yellow and greenish color shale and marl. Thin layered limestones are oolitic, especially in upper parts and are white to red in color, with fossils of Late Paleozoic age. It seems to be in relation to Sevan-Akera (Berberian and Amidi, 1981). As a whole, the Lahroud Complex is cut by two fault systems, the Angot Fault complex with a NW-SE trend, and the Alahyarlou Fault in the west with N-S trend (Fig. 2).



Figure3. Above photographs are from rock units in the Lahroud area. Post Paleozoic limestone has been overlain onto the ophiolitic assemblage, gabbro, ultramafic unit, pillow lavas, serpentinites. A) The gabbro is over-thrusted by a younger andesite to basaltic andesite unit (Cretaceous), B) Paleozoic Pelagic limestone, C) Pelagic limestone on gabbro and ultramafic units with partly listwanite patches in orange color, tectonized contact and with hydrothermal alteration affects. D) Limestones overthrusted on the serpentinites. Abbreviations: Pel. Lim.: Pelagic limestone, Lw.: Listwanite; Lim.: Limestone; Serp. Serpentinite.

Petrography

Intrusive rocks

Plutonic rocks of the Lahroud ophiolite are exposed in the area, are composed of mafic to ultramafic rock units, with ultramafic rocks comprising nearly 2/3 of the entire ophiolite complex. The crustal sequence consisted largely of gabbros which are present as relatively small bodies and appears to be fragmented parts of tectonically dismembered large units. For the purpose of the study and according to the geochemistry and metamorphic degree, it is divided into Meta-gabbro- and younger gabbro units. Meta-gabbro which is mainly cumulate-banded gabbro comprises porphyroclasts of plagioclase (60% to 70%), pyroxene (30% to 40%), and amphiboles (0 to 5%) with intergranular and interstitial textures (Fig. 4A). Some of the rocks show some degrees of alteration that are partly pervasive. This unit is also affected by metamorphic events (amphibole-muscovite schist). Amphibolite has been dated

(Fig. 4A, B) and the results are described in this research. This crustal sequence consists of gabbro that is presented as small body and appeared as fragments of tectonically larger units (Fig. 4C). Most of these rocks exhibit some degree of serpentinization, which is locally pervasive.

The younger gabbro in the area is fresher and is largely composed of non-cumulate rocks with ophitic textures. It has partly small outcrops in western part of the region, is formed in the sequence consists of porphyroclasts of plagioclase (60% to 70%), pyroxene (30% to 40%), and amphiboles (0 to 5%) with inter-granular or interstitial textures. Pyroxene and plagioclase phynocrysts are grained with ophitic texture (Fig. 4D). This gabbro is respectively fresher that the old gabbro except very thinly calcifications.

Extrusive rock units

Extrusive rocks in the Lahroud ophiolite occur as massive, coherent bodies and rarely as pillow lavas. Most of the basalts, andesite, basaltic andesite and pillow lavas are altered to some extent, but relict igneous textures are well preserved. In most areas, the pillow lavas are intercalated with a mixture of lava clasts, shallow- and deep-water fossiliferous limestone clasts, and reddish clasts of jasper and chert. These blocks of limestone and cherts vary in size from a few meters to tens of meters.

In hand specimen, these rocks are largely aphyric and contain microvesicles that are filled with zeolite and carbonate. Studies under microscope, show largely aphyric textures (Fig. 4 E). Andesite and basaltic andesite contains phenocrysts of plagioclase occurring alone or as clusters of closely packed crystals in a matrix of plagioclase microlites. The majority of these rocks are porphyritic, containing phenocrysts of plagioclase and hornblende and to lesser extent clinopyroxene in a fine-grained matrix. Basalts in the region are mostly spilitic at the surface.

In the upper part of the area, the pillow lavas are overlain by deep-water fossiliferous limestones. This block of limestone unit and cherts vary in thickness from a few meters to some ten meters, and interbedded with reddish jasper and cherts.

Metamorphism

Old rocks from mafic unites in the ophiolite sequence, are metamorphosed. Amphibolite, muscovite schist and sillimanite schists are major features which are observed in the region. Metamorphic rocks comprise schist with porphyroblasts of muscovite and sillimanite as major phenocrysts (Fig. 4F).

Mafic rocks are mostly converted to Serpentines. It is also has remarkably green and talky outcrops in the area (Fig. 4G, H).

Silica-carbonates (Listwanites) Listwanites form as a result of the chemical reaction between serpentinite and CO₂-rich fluids. These fluids usually migrate along faults or fractures along the contact of serpentinite and the adjacent country rocks in the region. Freshly broken listwanites have a green-orange color and weathered surface of listwanites outcropped as a gossanous box work texture and a brown-red color due to the preferential breakdown of the ferro-magnesium carbonates (Fig. 3C). Silica carbonates in the Lahroud region commonly found in silica-carbonates include chlorite, talc, residual serpentine without chromite, or mineralization bearing of any sulfides.



Figure4. Photomicrographs (cross polarized light) of metamorphic rocks; A) Amphibolite, the oldest rock with amphiboles that have been dated; B, C) meta-gabbro with phenocrysts of pyroxene and micas; D) freshest gabbro in the area as above part of the complex; E) basalt; spilite, F); Metamorphosed rocks, muscovite-sillimanite schist. G) Serpentinite with a large amount of talc; H) Serpentinite in hand specimen; I) Amphibole and muscovite in metamorphosed rock. Abbreviations: Amph: Amphibole; Talc: Talc; Plag: Plagioclase; Bio: Biotite; Pyrox: Pyroxene; Horn: Hornblende; OI: Olivine; Sil: Sillimanite; Mus: Muscovite; Serp: Serpentinite.

Analytical methods

ICP-MS, -OES

Seven samples were pulped for analysis of major and trace elements, including rare earth element (REE), by fusion inductively coupled plasma-mass spectroscopy and -optical emission spectroscopy (ICP-MS, -OES) at Amdel in Australia. The detection limit was 0.01% for all major element oxides and 0.5–1 ppm for rare earth elements. The representative major- and trace-element compositions of the Lahroud rocks are presented as below:

			-											
Sample No	SiO ₂	TiO ₂	Al ₂ O ₃	Fe ₂ O ₃ **	FeO**	MnO	MgO	CaO	Na ₂ O	K ₂ O	P2O5	LOI	Total	
Unit	wt.%	wt.%	wt.%	wt.%	wt.%	wt.%	wt.%	wt.%	wt.%	wt.%	wt.%	wt.%	wt.%	
Oph15	41.53	1.017	16.09	2.29	6.364	0.19	12.7	12.6	0.229	0.16	0.08	3.53	97.5	
Oph17	44.99	2.199	13.99	4.2	4.691	0.23	2.31	10.2	5.341	2.32	0.89	5.74	97.1	
Oph21	49.57	1.818	16.18	4.1	5.114	0.16	5.87	5.93	4.192	2.05	0.3	3.46	99.5	
Oph22	52.56	1.051	15.98	3.11	5.303	0.18	5.79	7.16	2.965	2.49	0.27	2.32	99.9	
Oph4	53.87	1.367	17.15	4.49	5.127	0.14	2.92	4.59	7.319	0.2	0.39	2.72	101.1	
Oph7	45.77	2.268	14.43	3.91	4.743	0.22	2.17	9.86	5.297	2.16	0.96	7.74	98	
Pi-BAS	55.32	1.317	15.65	3.5	5.48	0.17	4.44	4.74	3.397	1.62	0.26	4.28	100	
Plgr	77.08	0.083	12.5	0.83	0.664	0.01	0.12	0.13	3.572	4.42	0.009	0.62	100	
Sample No	Ba	Rb	Sr	Cs	Li	т	Nb	Hf	Zr	Y	Th	U	Cr	
Unit	ppm	ppm	ppm	ppm	ppm	ppm	ppm	ppm	ppm	ppm	ppm	ppm	ppm	
Oph15	38.8	4.8	423	0.6	105	< 0.1	1.5	0.7	24	24.4	0.3	< 0.1	50	
Oph17	282	41.7	398	0.9	12	0.1	49.1	4.3	329	39.7	5.9	1.4	<50	
Oph21	226	53.6	390	0.9	29	0.3	17.5	3.8	183	23.7	2.7	0.9	100	
Oph22	1280	72.8	500	1	19	0.3	10	0.6	17	20.2	3.7	0.9	100	
Oph4	46.7	2.8	215	<0.1	7	< 0.1	13	4	186	34.1	2.6	0.7	<50	
Oph7	424	39.4	387	1.2	13	0.1	47.5	5.1	248	40.2	6	1.6	<50	
Pi-BAS	158	36.6	209	1.1	18	0.2	7	3.5	137	32.9	1.8	0.5	<50	
Plgr	323	132	97	1	6.5	0.9	11.5	5	141	31.1	22.2	6.7	<50	
Sample No	Ni	Со	Sc	v	Cu	Pb	Zn	Bi	Cd	Sn	w	Мо	В	Be
Unit	ppm	ppm	ppm	ppm	ppm	ppm	ppm	ppm	ppm	ppm	ppm	ppm	ppm	ppm
Oph15	141	37.6	38.5	285	283	1	54	< 0.1	< 0.5	<1	0.5	< 0.5	<5	0.2
Oph17	11	19.2	11.8	133	22	5	144	< 0.1	<0.5	5	< 0.5	1	<5	0.7
Oph21	65	34.8	23	180	60	3	98	< 0.1	<0.5	<1	< 0.5	1	<5	1.3
Oph22	41	31.1	27.5	200	12	17	92	< 0.1	< 0.5	<1	< 0.5	< 0.5	<5	1.3
Oph4	<1	16	17	95	16	3	87	< 0.1	<0.5	1	< 0.5	0.5	<5	0.8
Oph7	7	23.1	12.5	120	15	8	120	< 0.1	<0.5	2	< 0.5	2	<5	0.9
Pi-BAS	16	24	29	160	40	4	103	< 0.1	< 0.5	1	< 0.5	0.5	<5	0.6
Plgr	6	1.6	2.5	<5	6	10	24	0.2	<0.5	3	2	4	<5	1.8
Sample No	Ag	Re	S	As	Sb	Te	La	Ce	Pr	Nd	Sm	Eu	Gd	Ть
Unit	ppm	ppm	ppm	ppm	ppm	ppm	ppm	ppm	ppm	ppm	ppm	ppm	ppm	ppm
Oph15	< 0.1	< 0.1	150	<1	< 0.2	<0.2	2.9	8.2	1.2	6.18	2.2	0.78	3	0.52
Oph17	<0.1	< 0.1	173	9	0.2	< 0.2	39.9	81.1	10.02	42.01	7.99	2.97	7.87	1.27

Table 2: Bulk rock geochemical analyses of selected samples from the Lahroud ophiolite

Oph21	< 0.1	< 0.1	200	2	0.2	<0.2	22.3	46.8	5.6	22.2	4.8	1.58	4.45	0.64
Oph22	<0.1	< 0.1	50	5	1	<0.2	24.3	46.1	6.2	23.1	4.78	1.6	4.6	0.64
Oph4	<0.1	< 0.1	<50	<1	< 0.2	<0.2	19.5	40.1	5.25	21.7	5.46	1.94	5.45	0.86
Oph7	<0.1	< 0.1	250	5	0.2	<0.2	40.6	79.1	9.75	38.4	8.58	3.12	8.05	1.22
Pi-BAS	< 0.1	< 0.1	50	4	< 0.2	< 0.2	12.7	26.4	3.85	16.9	4.86	1.66	5.3	0.86
Plgr	< 0.1	< 0.1	50	6	1.2	< 0.2	45	83.9	9.75	34.2	7.04	0.7	5.8	0.98
Sample No	Dy	Но	Er	Tm	Yb	Lu	Ge	Nb/Y	Zr/Y	La/Sm	Th/Nb	Nb/Yb	Th/Yb	
Unit	ppm	ррт	ppm	ppm	ppm	ppm	ppm	ppm	ррт	ppm	ppm	ppm	ppm	
Oph15	3.66	0.8	2.35	0.35	2.2	0.32	<0.5	0.06	0.98	1.31	0.2	0.682	0.14	
Oph15 Oph17	3.66 7.45	0.8 1.32	2.35 3.73	0.35 0.5	2.2 3.22	0.32	<0.5 <0.5	0.06 1.48	0.98 8.28	1.31 4.99	0.2 0.12	0.682 15.24	0.14	
Oph15 Oph17 Oph21	3.66 7.45 3.88	0.8 1.32 0.76	2.35 3.73 2.05	0.35 0.5 0.3	2.2 3.22 1.75	0.32 0.5 0.26	<0.5 <0.5 <0.5	0.06 1.48 0.73	0.98 8.28 7.72	1.31 4.99 4.64	0.2 0.12 0.15	0.682 15.24 10	0.14 1.83 1.54	
Oph15 Oph17 Oph21 Oph22	3.66 7.45 3.88 3.82	0.8 1.32 0.76 0.78	2.35 3.73 2.05 2.05	0.35 0.5 0.3 0.3	2.2 3.22 1.75 1.85	0.32 0.5 0.26 0.28	<0.5 <0.5 <0.5 <0.5	0.06 1.48 0.73 0.49	0.98 8.28 7.72 0.84	1.31 4.99 4.64 5.08	0.2 0.12 0.15 0.37	0.682 15.24 10 5.41	0.14 1.83 1.54 2	
Oph15 Oph17 Oph21 Oph22 Oph4	3.66 7.45 3.88 3.82 5.88	0.8 1.32 0.76 0.78 1.14	2.35 3.73 2.05 2.05 3.2	0.35 0.5 0.3 0.3 0.45	2.2 3.22 1.75 1.85 3.05	0.32 0.5 0.26 0.28 0.48	<0.5 <0.5 <0.5 <0.5 <0.5	0.06 1.48 0.73 0.49 0.38	0.98 8.28 7.72 0.84 5.45	1.31 4.99 4.64 5.08 3.57	0.2 0.12 0.15 0.37 0.2	0.682 15.24 10 5.41 4.3	0.14 1.83 1.54 2 0.85	
Oph15 Oph17 Oph21 Oph22 Oph4 Oph7	3.66 7.45 3.88 3.82 5.88 7.84	0.8 1.32 0.76 0.78 1.14 1.46	2.35 3.73 2.05 2.05 3.2 3.85	0.35 0.5 0.3 0.3 0.45 0.55	2.2 3.22 1.75 1.85 3.05 3.35	0.32 0.5 0.26 0.28 0.48 0.5	<0.5 <0.5 <0.5 <0.5 <0.5 <0.5	0.06 1.48 0.73 0.49 0.38 1.18	0.98 8.28 7.72 0.84 5.45 6.16	1.31 4.99 4.64 5.08 3.57 4.73	0.2 0.12 0.15 0.37 0.2 0.13	0.682 15.24 10 5.41 4.3 14.2	0.14 1.83 1.54 2 0.85 1.79	
Oph15 Oph17 Oph21 Oph22 Oph4 Oph7 Pi-BAS	3.66 7.45 3.88 3.82 5.88 7.84 6.02	0.8 1.32 0.76 0.78 1.14 1.46 1.16	2.35 3.73 2.05 2.05 3.2 3.85 3.35	0.35 0.5 0.3 0.3 0.45 0.55 0.45	2.2 3.22 1.75 1.85 3.05 3.35 3.05	0.32 0.5 0.26 0.28 0.48 0.5 0.46	<0.5 <0.5 <0.5 <0.5 <0.5 <0.5 <0.5 <0.5	0.06 1.48 0.73 0.49 0.38 1.18 0.21	0.98 8.28 7.72 0.84 5.45 6.16 4.16	1.31 4.99 4.64 5.08 3.57 4.73 2.61	0.2 0.12 0.15 0.37 0.2 0.13 0.26	0.682 15.24 10 5.41 4.3 14.2 2.3	0.14 1.83 1.54 2 0.85 1.79 0.59	

Ar-Ar Dating

Two samples of muscovite and glass minerals (about 40-50 grains per samples) were prepared from representative samples from basalt unit and metamorphosed rock in Lahroud, following crushing, washing, and handpicking under the binocular microscope.

Muscovite (200–300 μ m) from sample OPh15 representing the metamorphosed rock and concentrates from sample OPh7 also from the basalt unit was prepared. The freshest samples were used in this regard. The mineral separates were wrapped in aluminum foil and stacked in an irradiation capsule with similar-aged samples and neutron flux monitors (Fish Canyon Tuff sanidine (FCs), 28.02 Ma (Renne et al., 1998). The samples were irradiated at the McMaster Nuclear Reactor in Hamilton, Ontario for 43 MWH with a neutron flux of approximately 6x1013 neutrons/cm2/s. Analyses (n=30) of 6 neutron flux monitor positions produced errors of <0.5% in the J values.

The samples were analyzed at the Noble Gas Laboratory, Pacific Centre for Isotopic and Geochemical Research (PCIGR), University of British Columbia, Vancouver, BC, Canada. The mineral separates were step-heated at incrementally higher powers in the defocused beam of a 10W CO2 laser (New Wave Research MIR10) until fused. The gas evolved from each step was analyzed by a VG5400 mass spectrometer equipped with an ion-counting electron multiplier. All measurements were corrected for total system blank, mass spectrometer sensitivity, mass discrimination, radioactive decay during and subsequent to irradiation, as well as interfering Ar from atmospheric contamination and the irradiation of Ca, Cl and K (Isotope production ratios: $(40Ar/39Ar)K=0.0302\pm0.00006$, $(37Ar/39Ar)Ca=1416.4\pm0.5$, $(36Ar/39Ar)Ca=0.3952\pm0.0004$, Ca/K=1.83±0.01(37ArCa/39ArK).

In general, total gas ages, plateau ages, and inverse isochron ages are concordant within 2σ error, which rules out major problems caused by either excess argon or partial argon loss (Table 3).

Laser	Isotope Ratios												
	SH175 glass			(sample,	/mineral)								
Power(%)	40Ar/39Ar	$l\sigma$	37Ar/39Ar	$l\sigma$	36Ar/39Ar	$l\sigma$	Ca/K	Cl/K	%40Ar	f 39Ar	40Ar*/39ArK	Age	2σ
									um				
2.00 W	1196.37	17.80	9.89	0.35	4.121	0.107	18.24		101.71	0.31	20.646	-209.67	± 560.55
2.20 W	765.79	6.27	8.76	0.26	2.641	0.059	16.14		101.81	0.95	13.982	-139.30	± 341.21
2.30 W	275.92	5.76	19.26	0.72	0.941	0.028	35.75		100.16	0.51	0.439	-4.22	± 120.86
2.60 W	35.21	0.28	72.69	1.34	0.094	0.002	140.33		61.57	3.29	14.263	131.83	± 12.00
3.30 W	36.99	0.22	130.19	2.28	0.104	0.002	262.57		54.50	5.81	18.536	169.53	± 13.86
4.00 W	24.84	0.14	90.18	1.64	0.052	0.001	176.39		32.10	12.40	18.013	164.96	± 7.03
4.70 W	21.08	0.11	11.75	0.21	0.006	0.000	21.70		4.36	47.76	20.335	185.16	± 1.98
5.00 W	22.34	0.12	8.55	0.15	0.006	0.000	15.76		4.25	20.05	21.520	195.39	± 2.11
5.30 W	23.86	0.16	19.98	0.38	0.014	0.001	37.12		10.78	5.24	21.592	196.01	± 4.23
5.90 W	25.45	0.19	0.21	0.02	0.014	0.001	0.38		16.75	3.06	21.189	192.54	± 4.11
6.70 W	40.22	0.42	0.51	0.06	0.077	0.004	0.93		56.60	0.63	17.462	160.13	± 19.03
Power(%)	40Ar/39Ar	$l\sigma$	37Ar/39Ar	$l\sigma$	36Ar/39Ar	$l\sigma$	Ca/K	Cl∕K	%40Ar atm	f 39Ar	40Ar*/39ArK	1σ	
Total/Average	24.738	0.056	0.475	0.015	0.0071	0.0001				100.00	20.765	0.073	1
			Volumo		Integrated								
J = 0.005302	24 ± 0.0000265		39ArK =	0.495	Date =	188.97	± 1.26	Ма					
Plateau Age = no	ot defined	rogulta pl	ataau stans: Mada	1.1 Solutio	$(\pm 0.5\% \text{ conf})$. 11	$\Lambda = 10$	0.2 ± 6.5	Ma				
Inverse isochron	(contenation age)	results, pi	ateau steps: Mode	I I Solutio	n (±95% -com.) o	111	Age = 15	0.5 ± 0.5	wia				
points	271 . 20 . MOUD	20 D	1 1 1 1	00520242	20(2)		U						
40/36 intercept:	271 ± 29 MSWD	= 20, Pro	bability = 0 (at J=.	.00530242	±.3% 2s)		C						
40/36 intercept:	271 ± 29 MSWD	= 20, Pro	bability = 0 (at J=.	.00530242	±.3% 2s)								
40/36 intercept:	271 ± 29 MSWD Isotope Ratios SH179	= 20, Pro	bability = 0 (at J=.	.00530242 (sample	(mineral)								
points 40/36 intercept: Laser	271 ± 29 MSWD Isotope Ratios SH179 muscovite 404+2047	= 20, Pro	bability = 0 (at J=.	.00530242 (sample	±.3% 2s) /mineral) 264r/204r		C=/V	CL/K	0/404-	£204.r	404-*/204-1/		2-
points 40/36 intercept: Laser Power(%)	271 ± 29 MSWD Isotope Ratios SH179 muscovite 40Ar/39Ar	= 20, Pro	bability = 0 (at J=. 37Ar/39Ar	.00530242 (sample: 1σ	±.3% 2s) /mineral) 36Ar/39Ar	Ισ	Ca/K	СѴК	%40Ar atm	f 39Ar	40Ar*/39ArK	Age	2σ
points 40/36 intercept: Laser Power(%)	271 ± 29 MSWD Isotope Ratios SH179 muscovite 40Ar/39Ar	= 20, Pro	bability = 0 (at J=. 	.00530242 (sample 1σ	±.3% 2s) /mineral) 36Ar/39Ar	Ισ	Ca/K	СѴК	%40Ar atm	f 39Ar	40Ar*/39ArK	Age	2σ
points 40/36 intercept: Laser Power(%) 2.00 W	271 ± 29 MSWD Isotope Ratios SH179 muscovite 40Ar/39Ar 524.31	= 20, Pro <i>Ισ</i> 90.58	bability = 0 (at J=. 37Ar/39Ar 4.84	00530242 (sample 1σ 3.70	±.3% 2s) /mineral) 36Ar/39Ar 0.720	<i>1σ</i> 0.260	Ca/K 8.89	Cl/K	%40Ar atm 40.46	<i>f 39Ar</i> 0.06	<i>40Ar*/39ArK</i> 313.218	<i>Age</i> 1724.71	2σ 616.35
points 40/36 intercept: <i>Laser</i> <i>Power(%)</i> 2.00 W 2.20 W	271 ± 29 MSWD Isotope Ratios SH179 muscovite 40Ar/39Ar 524.31 109.49	= 20, Pro <i>Iσ</i> 90.58 3.56	bability = 0 (at J=. 37Ar/39Ar 4.84 10.72	.00530242 (sample 1σ 3.70 1.14	±.3% 2s) /mineral) 36Ar/39Ar 0.720 0.186	<i>Ισ</i> 0.260 0.045	<i>Ca/K</i> 8.89 19.78	CV/K	%40Ar atm 40.46 49.33	<i>f 39Ar</i> 0.06 0.32	<i>40Ar*/39ArK</i> 313.218 55.898	Age 1724.71 453.07	2σ ± 616.35 ± 192.83
points 40/36 intercept: <i>Laser</i> <i>Power(%)</i> 2.00 W 2.20 W 2.50 W	271 ± 29 MSWD Isotope Ratios SH179 muscovite 40Ar/39Ar 524.31 109.49 109.74	= 20, Pro <i>Iσ</i> 90.58 3.56 1.64	<i>37Ar/39Ar</i> 4.84 10.72 11.60	00530242 (sample 1σ 3.70 1.14 0.39	<pre>/mineral) //mineral) 36Ar/39Ar 0.720 0.186 0.251</pre>	<i>lσ</i> 0.260 0.045 0.007	Ca/K 8.89 19.78 21.43	СVК	%40Ar atm 40.46 49.33 66.83	<i>f 39Ar</i> 0.06 0.32 2.99	<i>40Ar*/39ArK</i> 313.218 55.898 36.707	Age 1724.71 453.07 309.96	2σ $^{\pm}_{616.35}$ $^{\pm}_{192.83}$ $^{\pm}_{32.81}$
points 40/36 intercept: <i>Laser</i> <i>Power(%)</i> 2.00 W 2.20 W 2.50 W 2.90 W	271 ± 29 MSWD Isotope Ratios SH179 muscovite 40Ar/39Ar 524.31 109.49 109.74 49.58	= 20, Pro <i>Iσ</i> 90.58 3.56 1.64 0.38	<i>37Ar/39Ar</i> 4.84 10.72 11.60 22.61	.00530242 (sample 1σ 3.70 1.14 0.39 0.42	(mineral) 36Ar/39Ar 0.720 0.186 0.251 0.048	<i>Ισ</i> 0.260 0.045 0.007 0.001	<i>Ca/K</i> 8.89 19.78 21.43 42.08	CVK	%40Ar atm 40.46 49.33 66.83 24.87	<i>f 39Ar</i> 0.06 0.32 2.99 26.69	<i>40Ar*/39ArK</i> 313.218 55.898 36.707 37.855	Age 1724.71 453.07 309.96 318.85	2σ $^{\pm}_{616.35}$ $^{\pm}_{192.83}$ $^{\pm}_{32.81}$ $^{\pm}_{6.93}$
points 40/36 intercept: <i>Laser</i> <i>Power(%)</i> 2.00 W 2.20 W 2.20 W 2.50 W 2.90 W 3.50 W	271 ± 29 MSWD Isotope Ratios SH179 muscovite 40Ar/39Ar 524.31 109.49 109.74 49.58 42.65	= 20, Pro <i>1σ</i> 90.58 3.56 1.64 0.38 0.25	<i>37Ar/39Ar</i> 4.84 10.72 11.60 22.61 8.49	.00530242 (sample 1σ 3.70 1.14 0.39 0.42 0.16	<pre>/*.3% 2s) //mineral) 36Ar/39Ar 0.720 0.186 0.251 0.048 0.007</pre>	<i>lσ</i> 0.260 0.045 0.007 0.001 0.000	Ca/K 8.89 19.78 21.43 42.08 15.64	СИК	%40Ar atm 40.46 49.33 66.83 24.87 3.09	<i>f 39Ar</i> 0.06 0.32 2.99 26.69 63.06	<i>40Ar*/39ArK</i> 313.218 55.898 36.707 37.855 41.578	Age 1724.71 453.07 309.96 318.85 347.38	2σ t = 616.35 t = 192.83 t = 32.81 t = 6.93 t = 3.87
points 40/36 intercept: <i>Laser</i> <i>Power(%)</i> 2.00 W 2.20 W 2.20 W 2.50 W 2.90 W 3.50 W 4.30 W	271 ± 29 MSWD Isotope Ratios SH179 muscovite 40Ar/39Ar 524.31 109.49 109.74 49.58 42.65 47.53	= 20, Pro <i>Iσ</i> 90.58 3.56 1.64 0.38 0.25 0.74	<i>37Ar/39Ar</i> 4.84 10.72 11.60 22.61 8.49 31.94	.00530242 (sample <i>Iσ</i> 3.70 1.14 0.39 0.42 0.16 0.72	(mineral) 36Ar/39Ar 0.720 0.186 0.251 0.048 0.007 0.035	<i>Ισ</i> 0.260 0.045 0.007 0.001 0.000 0.003	Ca/K 8.89 19.78 21.43 42.08 15.64 59.85	Cl/K	%40Ar atm 40.46 49.33 66.83 24.87 3.09 16.47	<i>f 39Ar</i> 0.06 0.32 2.99 26.69 63.06 4.51	<i>40Ar*/39ArK</i> 313.218 55.898 36.707 37.855 41.578 40.615	Age 1724.71 453.07 309.96 318.85 347.38 340.04	2σ $t^{\pm}_{616.35}$ $t^{\pm}_{192.83}$ $t^{\pm}_{32.81}$ $t^{\pm}_{6.93}$ $t^{\pm}_{3.87}$ $t^{\pm}_{16.24}$
points 40/36 intercept: <i>Laser</i> <i>Power(%)</i> 2.00 W 2.20 W 2.20 W 2.50 W 2.90 W 3.50 W 4.30 W 5.30 W	271 ± 29 MSWD Isotope Ratios SH179 muscovite 40Ar/39Ar 524.31 109.49 109.74 49.58 42.65 47.53 48.82	= 20, Pro <i>lσ</i> 90.58 3.56 1.64 0.38 0.25 0.74 0.60	<i>37Ar/39Ar</i> 4.84 10.72 11.60 22.61 8.49 31.94 32.91	.00530242 (sample]σ 3.70 1.14 0.39 0.42 0.16 0.72 1.16	<pre>////////////////////////////////////</pre>	<i>lσ</i> 0.260 0.045 0.007 0.001 0.000 0.003 0.005	Ca/K 8.89 19.78 21.43 42.08 15.64 59.85 61.71	СІ/К	%40Ar atm 40.46 49.33 66.83 24.87 3.09 16.47 3.39	<i>f 39Ar</i> 0.06 0.32 2.99 26.69 63.06 4.51 2.37	40Ar*/39ArK 313.218 55.898 36.707 37.855 41.578 40.615 48.293	Age 1724.71 453.07 309.96 318.85 347.38 340.04 397.71	2σ \pm 616.35 \pm 192.83 \pm 32.81 \pm 6.93 \pm 3.87 \pm 16.24 \pm 25.67
points 40/36 intercept: <i>Laser</i> <i>Power(%)</i> 2.00 W 2.20 W 2.20 W 2.50 W 2.50 W 2.90 W 3.50 W 4.30 W 5.30 W	271 ± 29 MSWD Isotope Ratios SH179 muscovite 40Ar/39Ar 524.31 109.49 109.74 49.58 42.65 47.53 48.82	 = 20, Pro 1σ 90.58 3.56 1.64 0.38 0.25 0.74 0.60 	bability = 0 (at J=. 37Ar/39Ar 4.84 10.72 11.60 22.61 8.49 31.94 32.91	.00530242 (sample <i>Iσ</i> 3.70 1.14 0.39 0.42 0.16 0.72 1.16	±.3% 2s) /mineral) 36Ar/39Ar 0.720 0.186 0.251 0.048 0.007 0.035 0.015	<i>Ισ</i> 0.260 0.045 0.007 0.001 0.000 0.003 0.005	Ca/K 8.89 19.78 21.43 42.08 15.64 59.85 61.71	CVK	%40Ar atm 40.46 49.33 66.83 24.87 3.09 16.47 3.39	<i>f 39Ar</i> 0.06 0.32 2.99 26.69 63.06 4.51 2.37	<i>40Ar*/39ArK</i> 313.218 55.898 36.707 37.855 41.578 40.615 48.293	<i>Age</i> 1724.71 453.07 309.96 318.85 347.38 340.04 397.71	2σ t^{\pm} $t^{616.35}$ t^{\pm} $t^{192.83}$ t^{\pm} 32.81 t^{\pm} 6.93 t^{\pm} 3.87 t^{\pm} 16.24 t^{\pm} 25.67
points 40/36 intercept: <i>Laser</i> <i>Power(%)</i> 2.00 W 2.20 W 2.20 W 2.20 W 2.50 W 2.90 W 3.50 W 4.30 W 5.30 W <i>Power(%)</i>	271 ± 29 MSWD Isotope Ratios SH179 muscovite 40Ar/39Ar 524.31 109.49 109.74 49.58 42.65 47.53 48.82 40Ar/39Ar	= 20, Pro <i>I</i> σ 90.58 3.56 1.64 0.38 0.25 0.74 0.60 <i>I</i> σ	<i>37Ar/39Ar</i> 4.84 10.72 11.60 22.61 8.49 31.94 32.91 <i>37Ar/39Ar</i>	(sample) (s	<pre>/mineral) 36Ar/39Ar 0.720 0.186 0.251 0.048 0.007 0.035 0.015 36Ar/39Ar</pre>	<i>lσ</i> 0.260 0.045 0.007 0.001 0.000 0.003 0.005 <i>lσ</i>	Ca/K 8.89 19.78 21.43 42.08 15.64 59.85 61.71 Ca/K	СѴК	%40Ar atm 40.46 49.33 66.83 24.87 3.09 16.47 3.39 %40Ar atm	<i>f 39Ar</i> 0.06 0.32 2.99 26.69 63.06 4.51 2.37 <i>f 39Ar</i>	40Ar*/39ArK 313.218 55.898 36.707 37.855 41.578 40.615 48.293 40Ar*/39ArK	Age 1724.71 453.07 309.96 318.85 347.38 340.04 397.71 <i>Ι</i> σ	2σ t = 616.35 t = 192.83 t = 32.81 t = 6.93 t = 3.87 t = 16.24 t = 25.67
points 40/36 intercept: <i>Laser</i> <i>Power(%)</i> 2.00 W 2.20 W 2.20 W 2.20 W 2.50 W 2.90 W 3.50 W 4.30 W 5.30 W <i>Power(%)</i> Total/Average	271 ± 29 MSWD Isotope Ratios SH179 muscovite 40Ar/39Ar 524.31 109.49 109.74 49.58 42.65 47.53 48.82 40Ar/39Ar 46.355	$= 20, Pro$ $l\sigma$ 90.58 3.56 1.64 0.38 0.25 0.74 0.60 $l\sigma$ 0.188	<i>37Ar/39Ar</i> 4.84 10.72 11.60 22.61 8.49 31.94 32.91 <i>37Ar/39Ar</i> 11.464	00530242 (sample]σ 3.70 1.14 0.39 0.42 0.16 0.72 1.16]σ 0.134	±.3% 2s) /mineral) 36Ar/39Ar 0.720 0.186 0.251 0.048 0.007 0.035 0.015 36Ar/39Ar 0.0088	<i>lσ</i> 0.260 0.045 0.007 0.001 0.000 0.003 0.005 <i>lσ</i> 0.0002	Ca/K 8.89 19.78 21.43 42.08 15.64 59.85 61.71 Ca/K	СІ/К	%40Ar atm 40.46 49.33 66.83 24.87 3.09 16.47 3.39 %40Ar atm	<i>f 39Ar</i> 0.06 0.32 2.99 26.69 63.06 4.51 2.37 <i>f 39Ar</i> 100.00	40Ar*/39ArK 313.218 55.898 36.707 37.855 41.578 40.615 48.293 40Ar*/39ArK 40.748	Age 1724.71 453.07 309.96 318.85 347.38 340.04 397.71 <i>Ισ</i> 0.214	2σ ± 616.35 ± 192.83 ± 32.81 ± 6.93 ± 3.87 ± 16.24 ± 25.67
points 40/36 intercept: <i>Laser</i> <i>Power(%)</i> 2.00 W 2.20 W 2.20 W 2.20 W 2.50 W 2.50 W 2.90 W 3.50 W 4.30 W 5.30 W <i>Power(%)</i> Total/Average	271 ± 29 MSWD Isotope Ratios SH179 muscovite 40Ar/39Ar 524.31 109.49 109.74 49.58 42.65 47.53 48.82 40Ar/39Ar 46.355 88 ± 0.000255	$= 20, Pro$ $I\sigma$ 90.58 3.56 1.64 0.38 0.25 0.74 0.60 $I\sigma$ 0.188	<i>37Ar/39Ar</i> 4.84 10.72 11.60 22.61 8.49 31.94 32.91 <i>37Ar/39Ar</i> 11.464 Volume	(sample) (s	 ±.3% 2s) /mineral) 36Ar/39Ar 0.720 0.186 0.251 0.048 0.007 0.035 0.015 36Ar/39Ar 0.0088 Integrated 	<i>lσ</i> 0.260 0.045 0.007 0.001 0.000 0.003 0.005 <i>lσ</i> 0.0002	Ca/K 8.89 19.78 21.43 42.08 15.64 59.85 61.71 Ca/K	CV/K	%40Ar atm 40.46 49.33 66.83 24.87 3.09 16.47 3.39 %40Ar atm	<i>f 39Ar</i> 0.06 0.32 2.99 26.69 63.06 4.51 2.37 <i>f 39Ar</i> 100.00	40Ar*/39ArK 313.218 55.898 36.707 37.855 41.578 40.615 48.293 40Ar*/39ArK 40.748	Age 1724.71 453.07 309.96 318.85 347.38 340.04 397.71 <i>Iσ</i> 0.214	2σ fit = 616.35 fit = 192.83 fit = 6.93 fit = 6.93 fit = 16.24 fit = 25.67
points 40/36 intercept: <i>Laser</i> <i>Power(%)</i> 2.00 W 2.20 W 2.20 W 2.20 W 2.50 W 2.90 W 3.50 W 4.30 W 5.30 W 5.30 W <i>Power(%)</i> Total/Average J = 0.005093 Plateau Age = 3-	271 ± 29 MSWD Isotope Ratios SH179 muscovite 40Ar/39Ar 524.31 109.49 109.74 49.58 42.65 47.53 48.82 40Ar/39Ar 46.355 38 ± 0.0000255 47.0 ± 4.9 Ma	$= 20, Pro$ $l\sigma$ 90.58 3.56 1.64 0.38 0.25 0.74 0.60 $l\sigma$ 0.188	<i>37Ar/39Ar</i> 4.84 10.72 11.60 22.61 8.49 31.94 32.91 <i>37Ar/39Ar</i> 11.464 Volume 39ArK = (2s, including J-	.00530242 (sample]σ 3.70 1.14 0.39 0.42 0.16 0.72 1.16]σ 0.134 0.052 -error of	±.3% 2s) /mineral) 36Ar/39Ar 0.720 0.186 0.251 0.048 0.007 0.035 0.015 36Ar/39Ar 0.0088 Integrated Date = MSWD = 0.77	<i>lσ</i> 0.260 0.045 0.007 0.001 0.003 0.003 0.005 <i>lσ</i> 0.0002 341.15	Ca/K 8.89 19.78 21.43 42.08 15.64 59.85 61.71 Ca/K ± 4.66	С//К С//К С//К	%40Ar atm 40.46 49.33 66.83 24.87 3.09 16.47 3.39 %40Ar atm	f 39Ar 0.06 0.32 2.99 26.69 63.06 4.51 2.37 f 39Ar 100.00 steps 5	40Ar*/39ArK 313.218 55.898 36.707 37.855 41.578 40.615 48.293 40Ar*/39ArK 40.748	Age 1724.71 453.07 309.96 318.85 347.38 340.04 397.71 $I\sigma$ 0.214	2σ ± 616.35 ± 192.83 ± 32.81 ± 6.93 ± 3.87 ± 16.24 ± 25.67
points 40/36 intercept: <i>Laser</i> <i>Power(%)</i> 2.00 W 2.20 W 2.20 W 2.20 W 2.50 W 2.90 W 3.50 W 4.30 W 5.30 W 5.30 W <i>Power(%)</i> Total/Average J = 0.005093 Plateau Age = 3- Inverse isochrom	271 ± 29 MSWD Isotope Ratios SH179 muscovite 40Ar/39Ar 524.31 109.49 109.74 49.58 42.65 47.53 48.82 40Ar/39Ar 46.355 38 ± 0.0000255 47.0 ± 4.9 Ma (correlation age)	$= 20, Pro$ $I\sigma$ 90.58 3.56 1.64 0.38 0.25 0.74 0.60 $I\sigma$ 0.188 results, pl	<i>37Ar/39Ar</i> <i>37Ar/39Ar</i> <i>4.84</i> 10.72 11.60 22.61 <i>8.49</i> 31.94 32.91 <i>37Ar/39Ar</i> <i>11.464</i> Volume <i>39ArK</i> = (2s, including J. <i>376</i> <i>376</i> <i>376</i> <i>396</i> <i>376</i> <i>396</i> <i>396</i> <i>396</i> <i>396</i> <i>396</i> <i>396</i> <i>396</i> <i>396</i> <i>396</i> <i>396</i> <i>396</i> <i>396</i> <i>396</i> <i>396</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i> <i>397</i>	.00530242 (sample 1σ 3.70 1.14 0.39 0.42 0.16 0.72 1.16 1σ 0.134 0.052 -error of 1 1 Solutio	$\frac{\pm .3\% 2s)}{/mineral)}$ $\frac{36Ar/39Ar}{0.720}$ 0.720 0.186 0.251 0.048 0.007 0.035 0.015 $\frac{36Ar/39Ar}{0.0088}$ Integrated Date = MSWD = 0.77 nn (±95% -conf.) on	<i>Iσ</i> 0.260 0.045 0.007 0.001 0.000 0.003 0.005 <i>Iσ</i> 0.0002 341.15 n 7 points	Ca/K 8.89 19.78 21.43 42.08 15.64 59.85 61.71 Ca/K ± 4.66 Age = 3 ²	CU/K CU/K CU/K Ma 67.6% o through 14 ± 18 M.	%40Ar atm 40.46 49.33 66.83 24.87 3.09 16.47 3.39 %40Ar atm	<i>f 39Ar</i> 0.06 0.32 2.99 26.69 63.06 4.51 2.37 <i>f 39Ar</i> 100.00 steps 5	40Ar*/39ArK 313.218 55.898 36.707 37.855 41.578 40.615 48.293 40Ar*/39ArK 40.748	Age 1724.71 453.07 309.96 318.85 347.38 340.04 397.71 <i>I</i> σ 0.214	2σ $t^{\pm}_{616.35}$ $t^{\pm}_{192.83}$ $t^{\pm}_{32.81}$ $t^{\pm}_{6.93}$ $t^{\pm}_{3.87}$ $t^{\pm}_{16.24}$ $t^{\pm}_{25.67}$

Table 3: Ar-Ar age data

Notes: Total fusion age, isochron age, inverse isochron age, and plateau age. All errors are 2σ .

Radiogenic Isotopes (Rb, Sr, Sm, Nd, Pb)

All rocks in the study area were subjected to moderately to intense alteration. It is impossible to find reasonably fresh igneous rocks in outcrop. Our studies were carried out on one sample whole-rock analyses of basaltic rock (pillow lava). This sample, in addition to six other samples, was analyzed litho-geochemically as noted above. A sample (basalt) as a whole rock after crushing and powdering, prepared for isotopic measurements followed the procedure at the University of British Columbia, Canada. Isotopic ratios of Nd, Pb, and Sr and isotope-dilution abundances of Nd, Sm, Pb, Th, U, Sr, and Rb were measured at the University of British Columbia, Pacific Centre for Isotopic and Geochemical Research (PCIGR) Vancouver, Canada (Table 4).

Sample Code	Oph7	Oph17		Oph7	Oph17
Lithology	Olivine Basalt	Basalt			•
Era	Cretaceous	Cretaceous			
⁸⁷ Sr/ ⁸⁶ Sr	0.705903	0.705901	²³⁸ U	0.002138919	0.001999832
⁸⁶ Sr/ ⁸⁸ Sr	0.1197	0.1121	²³⁸ U/ ²⁰⁴ Pb	4102327.399	4093214.399
¹⁴³ Nd/ ¹⁴⁴ Nd	0.512839	0.512783	²³⁵ U/ ²⁰⁴ Pb	29752.88221	29819.90123
¹⁴⁵ Nd/ ¹⁴⁴ Nd	0.348406	0.327351	²³² Th/ ²⁰⁴ Pb	2409825.989	2398998.885
¹⁴⁶ Nd/ ¹⁴⁴ Nd	0.7225	0.7223	ε ND	3.92	3.8
²⁰⁸ Pb/ ²⁰⁴ Pb	38.7763	38.7761	Sum Pb	74.04928081	74.03999872
²⁰⁷ Pb/ ²⁰⁴ Pb	15.6264	15.6258			
²⁰⁶ Pb/ ²⁰⁴ Pb	18.6465	18.6371			
⁸⁷ Rb/ ⁸⁶ Sr	0.6755	0.6749			
¹⁴⁴ Sm/ ¹⁴³ Nd	0.2915	0.2912			
⁸⁷ Rb/ ⁸⁶ Sr	79.8887	79.8858			
¹⁴⁷ Sm/ ¹⁴³ Nd	47419.261	46517.322			
²⁰⁶ Pb/ ²⁰⁴ Pb	18.6465	18.5899			
²⁰⁷ Pb/ ²⁰⁴ Pb	15.6264	15.5833			
²⁰⁸ Pb/ ²⁰⁴ Pb	38.7763	38.7659			

Table 4: Age corrected, Pb, Sm-Nd, Rb-Sr isotopic ratios for basalts and diabases from the Lahroud ophiolite

Geochemistry

Samples in the research have been analyzed using the methods that above mentioned. Eight rock samples from the Lahroud region were analyzed (see Table 2). The samples show a variable composition. The SiO₂ content of the Lahroud ophiolite ranges from 41 to 55 wt. %. The MG# [Mg/(Mg+Fe²⁺) molar ratio, assuming Fe²⁺=0.9(Fe) total given as Fe₂O₃ in Table 2] of the Lahroud samples are 0.8–4.5 wt.%. In addition, samples from the Lahroud ophiolite contains TiO₂ (0.08–1.3 wt. %). The granite show slow aluminum granite. Al₂O₃ compositions are 12.5 wt. %, but SiO₂ contents are high and almost 77 wt. %. However,

unlike other oceanic granites that show low contents of K_2O (4.4 wt. %); the sample from Lahroud has a high K₂O content. Pillow basalt from the complex shows a calc-alkaline affinity and contains less than 2% K₂O, which characterizes the basaltic samples as ophiolitic. Besides their moderate LOI ($\sim 0.06-5\%$), the basaltic rocks are more extensively altered, because of ocean floor hydrothermal activities and later metamorphism. Thereby, we used trace elements because of their relatively immobility via hydrothermal alteration processes and low grade metamorphic conditions (e.g., Ta, Hf, Ti, Nb, Y and Zr) in characterizing the basalts. Lahroud rock composition includes trachyandesite, andesite, basalt and andesitic basalt. Using the Zr/Y -Zr discriminant diagram, there is a clear distinction between basalts plotted within plate fields. Based on AFM plot all samples show calcalkaline series (Irvine and Baragar, 1971; Fig. 5A). Based on Meschede, 1986, the Zr-Nb-Y discrimination diagram illustrated that extrusive rocks plotted in four type fields. Andesite, andesite basalts and basalts show calc-alkaline chemical affinities (Fig. 5B). The pattern for basalts, based on chemical analysis, divided into two different groups; within plate alkaline basalt (AII) and within plate tholeiite and within plate tholeiites and volcanic arc basalts (C). We used discrimination diagrams (Wood 1980) to compare the geochemical characteristics for basalts in ternary diagrams like Th-Nb- Hf. The Lahroud basalts data plotted in D-type destructive plate margin basalt and is differentiated from the originated basaltic melt. Meanwhile, in the diagram (Th-Nb-Hf), all samples illustrate a slight depletion in Nb element content (Fig. 5C). A ratio of Nb/Y and Zr/Y for pillow lava samples ranges from 0.21 to 1.18 and 4.1 to 6.1, respectively (Table 5).

Table 5: Incom	patible trace ele	nent ratios in	crust and ma	antle reservoirs (from comp	oilations b	yу
	1				` 1		~

	Primitive Mantle	N-MORB	E-MORB	Continental Crust	HIMU-OIB	EMI-OIB	EMII	Lahroud
Zr/Nb	14.8	30		16.2	3.2-5	4.2-11.5	4.5-3	1.7-16
La/Nb	0.94	1.07		2.2	0.66-0.77	0.86-1.19	0.89-1.09	0.8-1.9
Ba/Nb	9	1.7-8	4.9-8.5	54	4.9-6.9	11.4-17.8	7.3-13.3	3.5-22
Ba/Th	77	6		124	49-77	103-154	67-84	14-129
Rb/Nb	0.91	0.36	205-230	4.7	0.35-0.38	0.88-1.17	0.59-0.85	0.2-7
K/Nb	323	210-350	0.06-0.08	1341	77-179	213-432	248-378	0.01-0.19
Th/Nb	0.117	0.025-0.071		0.44	0.078-0.101	0.105-0.122	0.111-0.157	0.2-0.25
Th/La	0.125	0.067		0.204	0.107-0.133	0.107-0.128	0.122-0.163	0.1-0.14
Ba/La	9.6	4		25	6.8-8.7	13.2-16.9	8.3-11.3	2.3-13.37



Figure5. A: AFM plot from Irvine and Baragar (1971), all samples plotted in the calc-alkaline area; B: Nb-Zr-Y geochemical discrimination diagram (Mechede, 1986) showing two types of basaltic rocks, AI, WPA (within plate alkaline basalts); WPT (within plate tholeiite); **B**, P-MORB (Mid ocean ridge basalt); C-D, VAB (Volcanic arc basalt); **C**: Th–Hf–Nb discrimination diagram (Wood, 1980) showing that the basalt data plot in the compositional field of (A) N-MORB type, (B) B- MORB type and tholeiitic basalt within plate and differentiates, (C) alkaline within plate basalt and within the plate basalts and differentiates, (D) destructive plate-margin basalts and differentiates. Samples from Lahroud fall in D area; destructive plate-margin basalts and differentiates. **D**: Chondrite-normalized incompatible trace element diagram for the rocks from the Lahroud ophiolite. Normalizing values for chondrite are from Sun and McDonough (1989). **E:** Chondrite-normalized REE patterns for (Oph 22, 21, 15) low- and high-level gabbros, (Oph 7, 4 and Pi-bas) basalt and diabases, (Plgr) granite. Normalizing values for chondrite are from Sun and McDonough (1989).

REE geochemistry

Table 2 lists REE results and Fig. 5 D, E show the chondrite-normalized REE patterns for samples of the entire sequence of rocks in the crustal section of the Lahroud ophiolite, representing the igneous stratigraphy of the ophiolite and the overall range of REE abundances. In Fig. 5D, E the REE patterns for rocks from gabbro and meta-gabbro units have been shown (Oph 22, 21, 15 are ultramafics; Oph 7, 4 and Pi-bas are basalts, and Plgr and granite rocks).

The gabbros have concentrations of REE (100-150 chondrite) and the patterns are characterized by enrichment in the light REE [(LaN/SmN)ave=0.81, where N means chondrite-normalized] and a very slight enrichment in the middle REE. Their primary mineralogy consists of hornblende and to a lesser extent plagioclase. The overall concentrations of MREE in the gabbros are slightly higher (~10_ chondrite); however, they have also smaller positive Eu anomalies [(Eu/Eu*)ave=1.32]. But for the only one metagabbro sample REE pattern characteristic has a flat feature with a slight enrichment in REE and with chondrite feature like. The REE patterns for the basalt samples are shown in Fig. 5E. The volcanic rocks are characterized by moderate REE enrichment [(LaN/SmN)ave=3.3] and positive Eu anomalies (Eu/ Eu*)ave=1.6]. The REE pattern for basalt samples is similar in shape to patterns of the rest of the intrusive samples (Figs. 5 E). The REE patterns of the basalt samples from Lahroud are distinctly different from the patterns for diabase samples from complex of the Mashhad ophiolites (Shafaii Moghadam, 2014), and from the sequence pillow lavas. The REE patterns of basalts from the Mashhad ophiolites are light-REE depleted. However, the Lahroud patterns resemble REE patterns of pillow and basalt samples from the Mashhad ophiolite of northeastern Iran (Shafaii Moghadam, 2014). All the patterns show strong LREE enrichment [(LaN/SmN)ave=3.47]. The REE patterns, together with clear similarities in the REE and Eu anomalies, may suggest that these rocks are cogenetic and were derived from the same parental melt. The REE patterns of the meta-gabbro and granite are shown in Fig. 5E. These patterns are parallel to those of the basalts and gabbro, with nearly identical LREE enrichment [(LaN/SmN)ave=1.59] and the same Eu pattern [(Eu/Eu*)ave=0.81]. However, the overall REE concentrations are higher, between 100 and 150_ the chondrite values

Isotope geochemistry

Three types of isotopic reservoirs are recognized in continental crust by Taylor et al. (1984) which were characterized in relation to Nd, Sr, and Pb isotopes. Zindler and Hart (1986) have detected five members in the compositions of mantle, that can state all observations on mid

ocean ridge and oceanic island basalts with a diversity of mixing process. The contents of the mantle sources are summarized in Table 4 and plotted as a series of generalized isotopic correlation diagrams (Fig. 6). Isotope data in pillow lavas and gabbros samples have been summarized in Fig. 6 and Table 4.

Based on Table 4, Lahroud samples Pb, Sr, and Nd contents show BSE to HIMU composition. Data for the Lahroud sample variation of the isotope ratio of Nd to Pb plot near to EM2, EM1 and indicate an enriched mantle source (adapted from Hart, 1986). The data for the Lahroud samples plot the data source in variation of the isotope ratio of Nd to Sr and plot the Bulk Silicate Earth (BSE) (Zindler and Hart, 1986). This composition is equivalent to a homogenous primitive mantle that formed by degassing of the planet. Some oceanic basalt has isotope composition closely linked to the compositions of the bulk earth to the present (Rollinson, 1993). Generally, the enrichment is likely to be related to a subduction environment; thereby crustal materials have been subducted into the mantle via subduction erosion or in a related process. EMII has an affinity of the upper continental crust which can represent a recycling of continentally raised sediments, continental crust, deformed ocean crust or oceanic-island crusts. Then it is suggested that the enrichment is related to the mixing of the sub continental lithosphere with the mantle. EMI with an affinity related to the lower crust, probably can represent a recycling process with lower crustal materials, and suggests enrichment by mantle metasomatism process. Weaver (1991) suggested that EMI, EMII are produced by mixing between HIMU mantle and subduction oceanic sediments (Rollinson, 1993; Plank and Langmuir, 1998).



Figure6. **A**. Variation of isotope ratio of Sr and Pb of the Lahroud basalts. (Adapted from Hart, 1986, Hawaiian data from Stille et al., 1983); **B**. Variation of isotope ratio of Nd and Sr of the Lahroud basalt. (Adapted from Hart, 1986); **C**. Isotopic correlation diagram ²⁰⁷Pb/²⁰⁴Pb vs. ²⁰⁶Pb/ ²⁰⁴Pb for separation of resource; DM: Depleted Mantle, BSE: bulk earth silicate, EMI, EMII: Enriched mantle, HIMU: Mantle with high U/Pb, PREMA: With high mantle composition. The Lahroud sample is located very close to EMII with ocean pelagic sedimentary. **D**. εNd vs. ²⁰⁶Pb/²⁰⁴Pb; **E**. ²⁰⁸Pb/²⁰⁴Pb vs. ²⁰⁶Pb/²⁰⁴Pb vs. ²⁰⁶Pb/

In Fig. 6C, the data for the Lahroud samples variation of the isotope ratios of ²⁰⁷Pb/²⁰⁴Pb, ²⁰⁶Pb/²⁰⁴Pb plot in near to EM2 and indicate enriched mantle source (adapted from Hart, 1986). Enriched mantle has variable ⁸⁷Sr/⁸⁶Sr; low ¹⁴³Nd/¹⁴⁴Nd, with high ²⁰⁷Pb/²⁰⁴Pb, and ²⁰⁸Pb/²⁰⁴Pb with given values of ²⁰⁶Pb/²⁰⁴Pb. Zindler and Hart (1986) differentiate between enriched mantle types II (EMII) with high ⁸⁷Sr/⁸⁶Sr (Fig. 6A, B). Because of similarities between these areas with Mashhad ophiolite, we plotted our isotope data with them for comparison. According to Fig. 6 C, D to F the Lahroud data plots on Pacific-Atlantic MORB and Indian MORB. And based on Fig. 6 D, E, F data show characteristics of Pacific Atlantic MORB and Indian MORB with subducted sediments. Shafaii Moghadam et al. (2014) presented evidence that Mashhad ophiolites have high and variable ²⁰⁷Pb/²⁰⁴Pb, ²⁰⁸Pb/²⁰⁴Pb, and ⁸⁷Sr/⁸⁶Sr, in magmatic rocks, that can reflect alteration and reflection of the rocks with seawater and/or pollution by continental crusts and (or) reflect evidence of subducted terrigenous sedimentary materials in the source.

⁴⁰Ar-³⁹Ar Dating

Despite of intensely weathering and alteration process in this old area, selected dating samples from a lot of rocks showed good results (some of selected minerals in samples didn't give suitable and reliable results). This 40 Ar/ 39 Ar incremental heating ages report is the first information of the cooling ages of the oceanic crust in the Lahroud ophiolite. We have presented all 40 Ar/ 39 Ar experimental results in Table 3, 6 and Fig. 7. In two samples, well plateaus are apparent, detecting accurate age of 347.0±4.9 Ma and 187.7±7.7 Ma for Amphibolite metamorphic unit and pillow lava, respectively.

Table6. ⁴⁰Ar-³⁹Ar measurements of muscovite separate in Amphibolite from the Lahroud ophiolite

complex	ζ
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Sample	Mineral	Normal Isochron	Inverse Isochron	⁴⁰ Ar/ ³⁶ Ar Plateau age
Oph15	Muscovite	343.0±13.0 Ma	344±18 Ma	347.0±4.9 Ma
Oph7	Glass	187.7±7.7 Ma	190.3±6.5 Ma	177.9±3.9 Ma

Two samples from mineral concentrates (40-50) grains were prepared from representative samples from pillow lava basaltic rock and Amphibolite in Lahroud, following crushing, washing, and handpicking under the binocular microscope. Glass concentrates from sample (Oph7) represent the Lahroud pillow lava cooling age and muscovite concentrates from sample Oph15 also from amphibolite metamorphic unit represent metamorphic age, were provided. The freshest sample per unit was used in this research. The samples were analyzed at the Noble Gas Laboratory, Pacific Centre for Isotopic and Geochemical Research, University of British Columbia, Vancouver, BC, Canada. The mineral separates were stepheated at incrementally higher powers in the defocused beam of a 10W. Generally, total gas ages, plateau ages, and inverse isochron ages are concordant within 2σ error, which rules out major problems caused by either excess argon or partial argon loss.



Figure 7. ⁴⁰Ar–³⁹Ar age spectra for moscovite (amphibolite) and glass (basalt) separated from the Lahroud ophiolite sequence.

Sample Oph7 produced a plateau age, and with isochron age revealed initial compositions near to the atmospheric value, to confirm a reliable total fusion age of 187.7 ± 7.7 Ma. The isochron age (190.3 ± 6.5 Ma) is concordant, but is slightly larger fit uncertainty than plateau age (Fig. 7B). The Lahroud ophiolite was mostly coeval with Mashhad and other Paleo-Tethys ophiolites in a W-E suture zone. But for the sample Oph15 which produce a plateau age, the isochron revealed an initial composition like the atmospheric values; and confirmed a

reliable plateau age (347 ± 4.9 Ma). The isochron age (343 ± 13 Ma) is concordant with it, but has a slightly larger fit uncertainty than plateau age (347 ± 4.9 Ma) (Fig. 7A).

Biostratigraphic studies

Thirteen pelagic samples from sedimentary units in the Lahroud ophiolite were provided for micro-paleontological study. Studies on thin-sections of productive samples have been illustrated in Fig. 8 (A to M) and are summarized in Table 7. Eftekharnezhad and Behroozi, 1991, have indicated Early Permian biostratigraphy age for radiolarites intercalations with turbiditic units and lavas through internal Iranian Paleo-Tethys suture zone with time. But Zanchetta et al., 2013 have indicated Late to Early Permian and Middle Permian time (~260 – 290 Ma). This is in accordance with ⁴⁰Ar/³⁹Ar dating of hornblende gabbros from Mashhad ophiolite with age of ca. 288 to 282 Ma (Ghazi et al., 2001), and U-Pb zircon date with 262.3±1 Ma for Anarak plagiogranite (Paleozoic ophiolites outcrop in CIM; and 380-382 Ma for Darreanjir, in Mashhad ophiolites, Shafaii Moghadam et al., 2014).

Table	e 7: L	ist of	tossils	found	ın pe	lagic	limestone	and	cherts	from	the	Lahroud	ophioli	te se	quence	e in
							NW	' Irai	ı							

Fossil Reference	Age	Fossils	Samples
Brady (1876)	Upper Mississippian?, Carboniferous	Nodosinella cf. concinna (Brady)	Oph2*
Brady (1876)	Upper Mississippian? to lower Pennsylvanian, Carboniferous	Nodosinella cf. cylindrical (Brady)	Oph3*
Hall (1858)	Mississippian?, Carboniferous	Endothyra baileyi (Hall)	Oph6*
Rozovskaya (1961)	Lower Permian to middle Permian	Endostaffella sp. (Rozovskaya)	Oph7*
Ozawa (1925)	Pennsylvanian, Carboniferous	Staffella molleri (Ozawa)	Oph11*
Miklukho-Maklay (1965)	Mississippian?; Carboniferous, central Asia	Bituberitina bicamerata (Miklukho-Maklay)	Oph12*
Cushman & Waters (1927)	Upper Pennsylvanian, Carboniferous	Paleotextularia grahamensis (Cushman and Waters)	Oph13*

*All fossils are studied by S. Senemari.



Figure8. A- Paleotextularia sp. (Upper Pennsylvanian, Carboniferous), B- a - Nodosinella cf. concinna Brady (Upper Mississippian?, Carboniferous), b – Pseudokahlerina sp. (Upper Carboniferous - Lower Permian), C- Nodosinella cf cylindrica (Brady) (Upper Mississippian? to Lower Pennsylvanian, Carboniferous); D- Earlandia sp. (Pennsylvanian, Carboniferous); E- Earlandia sp. (Pennsylvanian, Carboniferous); F- Nodosinella cf. concinna Brady (Upper Mississippian?, Carboniferous), G- Endostafella sp. (Artinskian to Kungurian, Permian); H- Laxoendothyra sp. (Tournaisian to Visean, Carboniferous); I- Pachyphloia sp. (Pennsylvanian, Carboniferous); K- a - a piece of Fusulinella (Staffella) sp. (Artinskian to Kungurian, Permian), b - Uralodiscus sp. (Visean, Carboniferous); L- Bituberitina sp. (Mississippian?, Carboniferous). (The taxa considered in the present figures and table 7 is referenced in Loeblich and Tappan (1988).

Discussion

This research aimed to demonstrate: 1- age constraint to understand "*petrogenesis*; 2comparison to other Paleozoic ophiolites in Asia; and 3- tectonic" settings. The ~160 Ma age difference for Carboniferous radiometric age of metamorphosed rocks and Carboniferous (Mississippian? to Pennsylvanian) bio-stratigraphy for other rock units of the Lahroud maficultramafic complex concludes this is a part of the ophiolite complex that formed in the Carboniferous.

Petrogenesis and tectonic implications: By way of comparison, we plotted data from other Paleozoic ophiolites in Iran against Lahroud volcanic and igneous rocks which show MORB characteristic with calc-alkaline affinity in this research. Trace element ratios are illustrated in Table 3 and Fig. 9, for comparing Lahroud with the northeastern ophiolite rocks in Iran. Shafaii Moghadam et al., 2014 showed that most rocks of Darrehanjir have high ratios of Th/Yb in contrast to Nb/Yb, and showed higher than mantle arrays, which indicates additional melt/fluids ratios from subduction related sediment or continental crust pollutions. Th/Yb ratios in Fariman Permian lavas which plotted within enriched MORB mantle arrays, suggested a few enrichments from mantle sources and a minor subduction related influences, similar to those produced by mantle plumes (Shafaii Moghadam et al., 2014) (Fig. 9A). Igneous rocks in Mashhad plotted between Darrehanjir (high Th/Nb) and Fariman (low Th/Nb) rocks and the Lahroud is close to the Mashhad ranges. They compared magmatic rocks from Mashhad ophiolites with Rasht, Jandagh-Anarak, and Aghdarband ophiolites that appeared in relation to subduction features; thereby our data sits very close to the Mashhad in the Nb/Y vs. Zr/Y ratio diagram (Fig. 9B).

Based on Faure and Mensing, 2004, Nd is mostly immobile at fluid environments unless ratio of water/rock reaches to $>10^5$ among alteration process, therefore, the ¹⁴³Nd/¹⁴⁴Nd in studied rocks may be not influenced with seawaters via alteration process. Mobility of U and Pb via hydrothermal alteration influenced Pb isotope contents in affected rocks (e.g. Faure and Mensing, 2004); U element additions by alteration can increase ²⁰⁶Pb/²⁰⁴Pb during the time. Due to immobility of Th in the alteration process and because of low abundances of ²³⁵U, 207 Pb/ 204 Pb and 208 Pb/ 204 Pb, these ratios in igneous rocks will reflect mantle sources (Liu et al., 2013). In detecting the contributions of sediment melts to a depleted mantle source in Darrehanjir-Mashhad magmatic rocks, Shafaii Moghadam et al. (2014) used diagram of ¹⁴³Nd/¹⁴⁴Nd Th/Nd vs. Th/Nd which was modeled by Liu et al. (2013; Fig. 9C). Because of the immobility of Nd and Th elements in a fluid that released from subduction related oceanic crust and with related sediment, whereas they are mobile in the resourced melt (Liu et al., 2013). According to Shafaii Moghadam et al. (2016), the samples show a linear trend in mantle wedges (depleted Indian MORB mantle) and sediments (with lesser radiogenic Pb in Izu-Bonin sediment; Plank et al., 2007) and the Lahroud sample illustrates a behavior similar to the Mashhad samples range (Fig. 9C). Mashhad and Fariman komatiites show high

¹⁴³Nd/¹⁴⁴Nd which suggests a highly depleted mantle source lacking in any subducted sediments. The complex of Devonian-Permian Darrehanjir-Mashhad that belonged to the Paleo-Tethys belt represented another Paleozoic oceanic basin that could be related to the CAOB (Central Asian Orogenic Belt; Shafaii Moghadam, et. al., 2016).



Figure 9. Trace element compositional variations of the Lahroud ophiolite. For comparison we plotted also data on the other Iranian Paleozic ophiolitic magmatic rocks. (A) Th/Yb vs. Nb/Yb, (B) Nb/Y vs. Zr/Y and Th/Nb vs. ¹⁴³Nd/¹⁴⁴Nd diagrams. Th/Yb vs. Nb/Yb diagram is from Pearce (2008) and Nb/Y vs. Zr/Y diagram from Fitton et al. (1997). Data on Oman ophiolites from Godard et al. (2003) and on Izu-Bonin-Mariana (IBM) forearc from Reagan et al., (2010). The Lahroud samples set in IBM fore arc basalt like Mashhad rocks. C- Plot of Th/Nd vs. ¹⁴³Nd/¹⁴⁴Nd for the Lahroud rock into other similar samples from Paleo-Tethys suturs zone samples. The Darrehanjir- Mashhad magmatic rocks are dispersed in this diagram, but in a binary array between mantle wedge and sediment components. The mantle wedge end-member is very depleted-Indian MORB mantle (DIMM; Stracke et al., 2003) with high¹⁴³Nd/¹⁴⁴Nd but low Th/Nd ratios (modified after Liu et al., 2013; Shafaii Moghadam et al., 2014).

Other Paleozoic ophiolites in Asia: The Tavric (Caucasus)-Kura (Turkey) mélange with some exotic block of Carboniferous-Permian pelagical limestones (Zonenshain et al., 1990; Sengor et al., 1988), suggests that Paleo-Tethys extended as far west as Turkey. The Trans-Caucasian massif, which is situated among the Great Caucasus in the north and the Lesser Caucasus in the south (Zakariadze et al., 2007), contained some evidence of Early Paleozoic metamorphosed ophiolite units containing sheeted dykes and pillow lava (Rolland et al., 2011; Gamkrelidze et al., 1999). A tectonized mélange zone, in the north of the Sevan-Akera suture zone includes allochthonous pieces from middle Paleozoic rocks and was affected by early Variscan metamorphosed rocks according to Ar-Ar dates (about 330 to 336 Ma; Treloar et al., 2009; Zakariadze et al., 2007). Other remnants of Paleo-Tethys are located in Turkmenistan (Boulin, 1988) that are interpreted as Early Paleozoic to Late Devonian deep ocean sedimentary rocks and ophiolite; in Kizilkaya and Tuar areas of Turkmenistan (Mirsakhanov, 1989); early Paleozoic- Late Devonian rocks in Afghanistan including flysch

type belonging to Cambrian and Ordovician to Silurian ages (Boulin, 1988); northeastern part of Afghanistan (in western part of Hindu Kush, Badakhshan areas; Boulin, 1988), and the Tibetan plateaus that also preserved Paleo-Tethys ocean remnants and following mixtures of the Laurasia and Gondwana super continents from the Paleozoic (Zhai et al., 2013; Sengor and Natalin, 1996). Furthermore to eastern parts, pillow lava from north-northeast Tadjikistan (pelagic limestone intercalation with Famennian-Tournaisian; (~360 Ma age). The Late Carboniferous and Early Permian intrusions with volcanism in the Mediterranean and Caucasus areas are related to the northward subduction trend of the Paleo-Tethys beneath Eurasia (Rolland et al., 2011; Stampfli & Borel, 2002). Low pressure and high temperature metamorphic units outcrop in the south of Georgia where Ar-Ar ages yield about 303 Ma (Carboniferous time from Rolland et al., 2011). High pressure rocks in the north of Iran (e.g. Shanderman in Rasht ophiolite; Omrani et al., 2013) with Ar-Ar age of 330 Ma; Zanchetta et al., 2009 are inferred as Paleo-Tethys subduction channel ways. Southward trend of subduction beneath the North of Tianshan arc (within CAOB; southeast of Kazakhstan) is probably responsible for the generation of ca. 343 Ma SSZ-type plagio-granites and E-MORB with OIB-like gabbros ca. 302 Ma ages (Li et al., 2014). The Carboniferous Kunlun arc located in the north of the Pamir Mountains and northwest of Tibet through to magmatism in Tadjikestan are related to northward vergent subduction settings (Schwab et al., 2004).

Iran Paleo-Tethys suture zone: The Paleo-Tethys opened in the Early Ordovician (Stampfli et al., 1991) and subduction was initiated ca. 380 Ma at the southern edge of Eurasia and was probably diachronous as Caucasus and Turkmenistan ophiolite which represents the Carboniferous SSZ-type Ocean crust. Zanchetta et al. (2013) concluded that the Paleozoic volcano-sedimentary units in Fariman- Darrehanjir areas are Permian. They also conclude that Fariman-Darrehanjir magmatic rocks developed as remnants of a magmatic arc in the south margin of Eurasia, in the upper part of the northward dipping Paleo-Tethys subduction zone that was an active margin from the Carboniferous. According to Shafaii Moghadam, et. al., 2014, U-Pb age data, geochemistry, and occurrences of abyssal peridotites in association with Devonian boninite to calc-alkaline gabbro/plagio-granites, supported their U-Pb age data and demonstrates a pulse of SSZ type ocean crust forming in the Early to Late Devonian (Fig. 10). Permian igneous rocks in association with terrigenous/pelagic sedimentary rocks could be representative of magmatism and sediments in the subduction related setting at the foot of the Eurasian margin, similar to the accretionary orogeny with fragments, in Devonian ocean crust and then followed by Permian magmatic rocks. Geochronology data for Granitegranodiorite intrusions into the Mashhad and Torbat-e-Jam ophiolites yielded U-Pb ages of 215 ± 4 and 217 ± 4 Ma (Karimpour et al., 2010) and 199.8 ± 3.7 and 217 ± 4 Ma (Zanchetta et al., 2013; Mirnejad et al., 2013). Northward subduction of the Paleo-Tethys oceanic lithosphere zone started at the southern margin of Eurasia in the Middle- Later Paleozoic (Zanchi et al., 2009; Metcalfe, 2006; Alavi, 1991;; Fig. 10) in Iran, Afghanistan, Caucasus and Turkey. The Lahroud is therefore inferred to represent a part of the Paleo-Tethys suture zone in the northwestern part of Iran, with a metamorphic event at 347 Ma as a part of metamorphic regime in a thickened oceanic crust (Fig. 7A) affected by subduction processing. Age dating from pillow lava (basalts) gives an age for closure of the ophiolitic system in the region (Fig. 7B).



Figure 10. Simplified chart showing the ages of magmatic and sedimentary sequences that constrain ages of Iranian Paleozoic ophiolites; greenish area inside dashed lines encompasses ophiolite ages. U-Pb dating of granitic pebble from Qara-Gheitan Formation (K) from Karimpour et al., (2011), Ar-Ar age data on the Mashhad ophiolites from Ghazi et al., (2001), U-Pb zircon ages on the Mashhad granites from Karimpour et al., (2010) and Mirnejad et al., (2013), Ar-Ar ages on the Rasht-Shanderman eclogites from Zanchetta et al., (2004), U-Pb zircon age data on the Anarak trondhjemites and Jandagh granites from Bagheri and Stampfli, (2008), U-Pb zircon data on the Torbatejam granites, granitic pebble from Qara-Gheitan Formation (Z) and granitic pebbles from Kashafrud Formation from Zanchetta et al., (2013). Ar-Ar ages on the Caucausus accretionary prism and high-T metamorphic rocks a from Rolland et al., (2011); data on the Turkmenistan and

Afghanistan pelagic sediments are from Garzanti and Gaetani (2002), Zakariadze et al. (2007), Treloar et al. (2009) and Boulin (1988) respectively.at 347-187 Ma (Modified after Shafaii Moghadam et al., 2014).

Conclusion

Our finding in this research suggested the followings: (1) The Lahroud ophiolite is apparently an ophiolitic complex in association with a basal metamorphic "zone" in Paleo-Tethys oceans; (2) The Lahroud ophiolite comprises a typical sequence of ophiolitic rocks; (3) the volcanic rocks represent the upper part of the ophiolite rock in the region, overlain by fossiliferous pelagic limestones and radiolarian cherts. Radiometric dating on amphibolite and basalt (pillow lava) clearly suggests the end of the Paleozoic (Carboniferous) as a maximum age for the beginning of formation of the ophiolitic complex. Chrono-stratigraphy data that resulted from pelagic limestone biostratigraphy studies yield the same age. These ages are (late Paleozoic) Mississippian? to Pennsylvanian. Younger age for pillow lava indicates continuation of "Lahroud oceanic crust" formation in the late Carboniferous to Early Jurassic. Consequently, it is inferred that the Lahroud Ophilite formed as a result of the closure of this part of the paleo-Tethys. We propose that subduction and closing of the oceanic crust started almost 340 Ma and we see a strong metamorphism as a finger print in the region. The scenario continued until 180 Ma and finished with basaltic pillow lava volcanic extrusive rocks formation. The results revealed that the Lahroud "ophiolite" is an igneous complex; biostratigraphy and Ar-Ar ages studies in the area indicated Permian age but the Ar-Ar ages of pillow lavas indicated Early Jurassic (PLIENSBACHIAN) crystallization ages but Early Carboniferous (TOURNAISIAN) for metamorphosed units. This is another piece in the jigsaw of the paleo-Tethys in northwestern Iran and its continuation to the Mashhad- Rasht suture zone. This research obtained additional constraints to the configuration of the Paleo-Tethys in the Devonian to Early Jurassic in Iran especially in the Lahroud region.

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